



## Optical properties of humic-like substances (HULIS) in biomass-burning aerosols

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# Optical properties of humic-like substances (HULIS) in biomass-burning aerosols

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## Abstract

We present here the optical properties of humic-like substances (HULIS) isolated from the fine fraction of biomass burning aerosol collected in the Amazon basin during the LBA-SMOCC (Large scale Biosphere atmosphere experiment in Amazonia – SMOke aerosols, Clouds, rainfall and Climate) experiment in September 2002. From the isolated HULIS, aerosol particles were generated and their scattering and absorption coefficients measured. The size distribution and mass of the particles were also recorded. The value of the index of refraction was derived from “closure” calculations based on particle size, scattering and absorption measurements. On average, the complex index of refraction at 532 nm of HULIS collected during day and nighttime was  $1.65-0.00187i$  and  $1.69-0.00163i$ , respectively. In addition, the imaginary part of the complex index of refraction was calculated using the measured absorption coefficient of the bulk HULIS.

The mass absorption coefficient of the HULIS was found to be quite low at 532 nm ( $0.031$  and  $0.029 \text{ m}^2\text{g}^{-1}$  for the day and night samples, respectively). However, due to the high Ångström exponent of HULIS (6–7) the specific absorption increases substantially towards shorter wavelengths ( $\sim 2-3 \text{ m}^2\text{g}^{-1}$  at 300 nm), causing a relatively high (up to 50%) contribution to the absorption at this wavelength. For the relative contribution of HULIS to light absorption in the entire solar spectrum, lower values (6.4–8.6%) are obtained, but those are still not negligible.

## 1. Introduction

During large-scale biomass burning events, vast amounts of aerosol are emitted directly into the atmosphere. The carbonaceous components of this aerosol are of utmost importance, since they make the dominant part of the particulate mass (PM) (Artaxo et al., 2002; Decesari et al., 2005). These aerosols scatter and absorb the incident light affecting the energy transfer through the atmosphere, and thereby the Earth’s energy budget and climate. During the burning process, depending on the oxygen supply,

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black carbon (BC) can be formed. Generally higher amounts of BC are emitted from flaming fires than from smoldering fires. Beside the soot (or BC), high amounts of carbonaceous compounds (including brown carbon or HULIS) can be found in the aerosol from biomass burning. Since HULIS are known to absorb light in the UV and visible range (Havers, 1998), the study of the absorption properties of this fraction is particularly important. We have some information about the absorption characteristics of the HULIS measured in continental urban and background aerosols (Havers, 1998), but their importance in the radiation transfer and their index of refraction, which is a fundamental parameter, are not very well known. Recently Kirchstetter et al. (2004) reported the optical parameters of the organic carbon from biomass burning aerosol using a filter-based method.

Since the specific absorption cross section of BC is the highest among the carbonaceous constituents of the aerosols, its contribution to the atmospheric absorption by aerosols is considerable. The brown carbon (including HULIS) is expected to be a weaker absorber, however, due to its higher abundance, its contribution to the absorption may also be important. In this study we present experimentally determined optical properties of humic-like substances isolated from biomass burning aerosol.

## 2. Method

### 2.1. Sampling, collection and preparation

The samples were collected at a tropical pasture site in Rondônia, Brazil, during the LBA-SMOCC campaign in September 2002, which corresponds to the most active biomass burning period in this region. Information about the sampling site can be found elsewhere (Andreae et al., 2002). Sampling was performed with a dichotomous virtual impactor (Solomon et al., 1983) mounted on a 10 m high tower. The samples were collected on Pallflex quartz filters, preheated at 600°C for at least 10 h. The air was sampled with a total flow of 16.33 m<sup>3</sup> h<sup>-1</sup>, and split inside the instrument at a 9:1 ratio

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in order to separate fine and coarse particles, respectively, with a 50% aerodynamic diameter ( $D_p$ ) cutpoint of  $2.5 \mu\text{m}$  (Solomon et al., 1983). Aerosols were collected separately during day and night, so that each filter corresponds to either daytime or nighttime aerosol to allow some investigation of diel effects. The typical sampling time for a given filter was about 10 h during daytime and about 12 h during nighttime.

The HULIS were isolated with the one-step procedure described by Varga et al. (2001) from the filters containing the fine fraction of the aerosol ( $D_p < 2.5 \mu\text{m}$ ). First, the water-soluble aerosol components were extracted by placing the filters for 24 h in MilliQ water. Then the extract was filtered through a Millipore membrane filter of  $0.45 \mu\text{m}$  pore size and the pH was adjusted to pH=2 with hydrochloric acid. The separation of HULIS from other dissolved components was performed on an Oasis HLB (1 g) column conditioned with methanol and activated with hydrochloric acid prior to the application of the samples. The HULIS fraction was then eluted with methanol and dried at room temperature. In order to have sufficient HULIS mass for subsequent measurements, all day and night sample extracts were combined into one daytime and nighttime sample, respectively.

## 2.2. Measurement setup and procedure

Figure 1 shows the measurement setup for determination of the optical parameters of the isolated HULIS. For particle generation by a TSI 3076 atomizer, the HULIS extract was redissolved in acetonitrile ( $\sim 0.3 \text{ g/l}$ ) containing 7% MilliQ water to ensure complete dissolution. Then the particles were dried ( $RH \sim 26\%$ ) and diluted by mixing the atomizer output flow with dry, filtered air. Subsequently, the aerosol was neutralized by a bipolar charger ( $^{65}\text{Ni}$ ) and passed through an impactor ( $1 \mu\text{m}$  cutoff diameter) as required for measuring the size distribution of submicron particles with a scanning mobility particle sizer (SMPS). Here we used a Grimm SMPS (Model 5.400). The generated HULIS particles were predominantly in the size range between 50 and 100 nm. For this size range, the performance deterioration of the SMPS due to particle diffusion is negligible and the sizing accuracy is about 3% (Kinney et al., 1991). For spherical

particles, the uncertainty in particle number concentration is about 10% mainly due to the relatively large uncertainties in particle charge distribution.

5 The particulate mass was recorded by a Tapered Element Oscillating Microbalance (TEOM, Rupprecht and Patashnick, Model 1400a), which was operated at a filter temperature of 30°C. We estimate the experimental uncertainty of the TEOM as few per-

cent.

10 The scattering coefficient was measured at a wavelength of 545 nm by an integrating nephelometer (Radiance Research, Model: M903) with an integrated temperature and *RH* sensor. Since this device is non-destructive and has negligible loss of submicron particles (less than 3%), it was operated upstream of the other aerosol devices. The nephelometer was calibrated with CO<sub>2</sub> (Anderson and Ogren, 1998) and the systematic error due to angular truncation of the viewing volume in the near forward (0 to 9°) and backward directions (170 to 180°) were corrected based on Mie calculations. The absorption coefficient was determined at 532 nm by a photoacoustic spectrometer (PAS).

15 Since the operating wavelengths of the PAS and nephelometer are very close to each other (within 13 nm), we neglected the wavelength dependence of the measured parameters within the 13 nm range, therefore we referred our measurements to 532 nm. The uncertainty of the absorption coefficient measured by the PAS is 5%. This was confirmed by calibration with NO<sub>2</sub> gas (Arnott et al., 2000) and for soot and biomass

20 burning aerosol by comparison with an optical extinction instrument (Schnaiter et al., 2005).

The experiments were carried out at *RH*~26%. One should recall that the isolated HULIS was redissolved in an acetonitrile-water mixture of 13:1. It is critical to make sure that all solvents volatilize from the particles under these conditions. Therefore, we

25 checked whether the particles were indeed dry in the measurement unit by measuring the size distribution of the generated HULIS particles with and without drying the sheath air in the SMPS system. The calculated values of the densities were basically the same. If the particles had contained solvents, which would have been lost in the dry sheath air, the densities of the particles would have had to be different. Furthermore,

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according to our best knowledge, acetonitrile and water does not make an azeotropic mixture, which would affect the drying process.

In addition to these on-line experiments, the absorption spectra of aqueous HULIS solutions were recorded by a Shimadzu UV-160A spectrophotometer from 300 to 700 nm in a 4 cm cell against water.

2.3. Calculation of refractive index

The complex index of refraction is derived from Mie “closure” calculations based on particle size, scattering and absorption measurements. The Mie code used was a modified version of that described by Guyon et al. (2003). Using the measured particle size distribution, the Mie code first takes the refractive index of water (1.33–0i) and calculates the corresponding scattering and absorption coefficients. These values are then compared to the observed scattering and absorption coefficients and the imaginary and real part of the refractive index are then increased stepwise until the calculated and measured optical coefficients agree to within 0.5%. The only noteworthy difference to the original code described by Guyon et al. (2003) is that in the present study the size distribution is fixed (determined by the SMPS) and does not need to be re-adjusted with refractive index as required for the optically determined size distribution used by Guyon et al. (2003).

In order to obtain information about the morphology of the generated aerosol, we collected particles on gold coated Nuclepore filters for SEM analysis. As shown in Fig. 2, the particles are indeed spherical and thus the application of the Mie code is justified.

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### 3. Results

#### 3.1. Absorption properties of HULIS

Table 1 summarizes the experimentally determined optical and physical properties of the HULIS isolated from pyrogenic fine aerosol. It can be seen from these data that there is no diel variation in either the real or the imaginary part of the refractive index of the HULIS.

In the literature the imaginary part of the refractive index of BC was found to vary from  $-0.45$  to  $-1$  (Reid et al., 2005 and references therein). However, it is important to note that for biomass burning aerosol a significantly lower imaginary part is obtained than for pure BC, indicating that soot represents only a minor fraction relative to organic compounds in pyrogenic aerosol. Recently, Guyon et al. (2003) reported values of  $1.41-0.013i$  and  $1.46-0.016i$  for the complex refractive index for ambient aerosols affected by biomass burning. Based on field measurements, D. Chand (personal communication, 2005) obtained  $1.47-1.53$  for the real part and  $0.01-0.02i$  for the imaginary part of the refractive index of the very same dry biomass burning aerosol that we used for our experiments. It should be kept in mind that their values refer to the entire fine aerosol including its black carbon fraction, whereas our data refer only to the HULIS fraction, a specific subset, which contains most of the light-absorbing organic compounds (Havers, 1998).

For one nighttime sample we also measured the spectrally resolved imaginary part of the index of refraction of the HULIS with a spectrophotometric method. We recorded the absorption spectra of the HULIS redissolved in water in a 4-cm cell, in order to minimize the contribution of baseline instabilities to the measurements at 532 nm. Considering the HULIS solution as a bulk material the index of refraction can be calculated using the equation (Bohren and Huffman, 1983; Gelencsér, 2004):  $-1/m(m) = \alpha\lambda/4\pi$  where  $\alpha$  is the absorption coefficient ( $\text{m}^{-1}$ ) and  $\lambda$  is the wavelength (m). Assuming that the index of refraction is volume additive and the bulk density of HULIS is  $1.5 \text{ g cm}^{-3}$ , the imaginary part of the index of refraction at 532 nm was calculated to be

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$-0.00184 \pm 0.0002$  in the concentration range of  $2.37 \times 10^{-2}$  and  $7.11 \times 10^{-3} \text{ g l}^{-1}$  HULIS.

In order to validate our on-line method described in Fig. 1, we calculated the scattering coefficient of pure, dry ammonium sulfate for the measured size distribution (using the index of refraction given in the literature) and compared with the measured scattering coefficient. We have found good agreement between the calculated and measured values ( $<10\%$  difference). Furthermore we measured the index of refraction for water-soluble nigrosin (Aldrich) both with our spectrophotometric (imaginary part only) and on-line method. The difference in the real part of the refractive index between the on-line and literature value was about 4%. With the spectrophotometric method we obtained almost the same imaginary part ( $-0.255$ ) reported in the literature ( $-0.26$ ) by Pinnick and Auvermann (1979). But the imaginary refractive index of nigrosin measured with the on-line method ( $-0.1795$ ) was different from that measured in aqueous solution. One explanation for this discrepancy might be some water uptake by the nigrosin particles (Pinnick, personal communication).

Since there were no diel variations in the imaginary part of the refractive index of HULIS particles, it is obvious that the same holds for the calculated mass absorption coefficients. Its value is low at 532 nm, at least relative to those of BC reported in the literature (e.g. Fuller et al., 1999). One should recall, however, that the mass absorption coefficient obtained for biomass burning fine aerosol is not high either (compared to that of the BC), the average value being  $0.37 \text{ m}^2 \text{ g}^{-1}$  at 532 nm (Chand et al., 2005). Secondly, the mass absorption coefficient of the HULIS increases sharply towards shorter wavelengths as indicated in Fig. 3 based on the spectrophotometric measurements.

In our experiments, the maximum of the number size distribution of the particles was at about 50–100 nm (diameter). Since for slightly absorbing particles the mass absorption coefficient is independent of particle size below 100 nm, increases by a factor of  $\sim 2$  in the range between 100 and 500 nm, and just slightly decreases again for 1–2  $\mu\text{m}$  particles (at a wavelength of 550 nm) (Horvath, 1993; Bohren and Huffman, 1983), our value may also be applied as a lower limit for absorbing organic aerosol (HULIS) present in ambient fine aerosol.

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The relatively high value of the mass scattering coefficients imply that HULIS efficiently scatter light in the visible range. The high value of the single scattering albedo obtained in our measurements (0.98) confirms this finding. Contrary to the absorption properties, the scattering properties seem to be different in HULIS isolated from day and night aerosol samples. It is known that HULIS may both be a primary aerosol component and be produced in photochemical processes, and the properties of these two components may well be different. The average mass scattering coefficient of biomass burning aerosol from the same sampling site in 2002 September was reported to be  $5.3 \text{ m}^2 \text{ g}^{-1}$  (Chand et al., 2005). Note that the density of the HULIS particles, an important parameter in calculations dealing with the role of HULIS in cloud formation, does not show diel variation.

### 3.2. Contribution of HULIS to light absorption in the visible range

Figure 3 shows the absorption spectra of HULIS isolated from day and night aerosol samples. Fitting power-law functions to the measured spectra reveals that the Ångström exponents are similar, and that in both cases the absorption is overestimated by the power-law relationship below about 350 nm. The Ångström exponent ( $\text{\AA}$ ) can be calculated for any range using the following equation (Reid et al, 2005):

$$\text{\AA}_{\lambda_1, \lambda_2} = -\ln(A_1/A_2)/\ln(\lambda_1/\lambda_2),$$

where  $A_1$  and  $A_2$  are the absorbances measured at wavelengths  $\lambda_1$  and  $\lambda_2$ , respectively. Using this equation we obtain 6.4 and 6.8 for the Ångström exponents of HULIS in day and night samples, respectively, in the wavelength range of 300–700 nm. Since  $\text{\AA}$  has a marked influence on the calculated absorption properties at shorter wavelengths, in our calculations we used two extreme values: 6 and 7. Important to note that Kirchstetter et al. (2004) obtained similar results (5.9) for the acetone soluble organic carbon in the range between 350 and 650 nm.

In this section the relative contribution of HULIS to light absorption in the visible will be estimated for typical conditions observed in the biomass burning period in Rondônia

during the SMOCC campaign. For the calculations, the absorption parameters of HULIS isolated from daytime samples are used. Based on our measurements, the absorption coefficient of HULIS is  $0.031 \text{ m}^2 \text{ g}^{-1}$  at 532 nm (see Table 1). Using the power-law relationship, the mass absorption coefficient of HULIS can be estimated for any wavelength. Since there is enhanced wavelength dependence in aerosol phase compared to the bulk phase (infinite large diameter), our results can be considered as a lower bound (Bond, 2001). According to our preliminary measurements we assumed a 1.9 organic mass-to-carbon ratio for HULIS in order to calculate the HULIS-contribution to the absorption on a carbon mass basis. The same mass-to-carbon ratio was found for the same fraction of the WSOC from a rural site (Kiss et al., 2002). We also assumed that the mass-to-carbon ratio of BC equals to 1.

Figure 4 shows the wavelength-dependence of the mass absorption coefficient of BC and that of the HULIS between 300 and 700 nm, using lower and upper estimates of the Ångström exponents of 6 and 7 as realistic lower and upper bounds, respectively. The mass absorption coefficient of soot was taken from the literature using the data reported by (Schnaiter et al., 2003). We used the value of 1.056 derived from the data in this article for the Ångström exponent of the diesel soot. We assumed that the average fine TC concentration during the biomass-burning dominated part of the LBA-SMOCC campaign is  $47.6 \mu\text{g m}^{-3}$  (Hoffer et al., 2005), and that 10% of the TC is BC with the same optical properties as the soot reported by Schnaiter et al. (2003). Furthermore, based on our preliminary chemical analysis of biomass burning aerosols from the LBA-SMOCC campaign we assumed that 35% by mass of TC represents the HULIS carbon fraction. The relative contribution of HULIS-carbon to the absorption can now be calculated using the mass absorption coefficients and the mass concentrations of soot and HULIS.

As we can see from Fig. 5, the contribution of HULIS to light absorption is only a few percent at 532 nm and continuously decreases towards longer wavelengths, but near the photochemically most active range, the absorption of organic compounds becomes significant relative to that of BC. For instance, at 300 nm, 35–50% of absorption may

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be attributed to HULIS. It should be noted that HULIS may be formed in photochemical reactions (Gelencsér et al., 2003; Hoffer et al., 2005), and since light-absorbing HULIS may also initiate the formation of OH and HO<sub>2</sub> radicals (Anastasio et al., 1994) a positive feedback mechanism is set up which could make the process particularly effective in the troposphere.

When integrating the contribution of HULIS to light absorption across the solar spectrum to assess their significance for the atmospheric radiation budget, lower total contributions are obtained. For a scale height of 3000 m, assuming the same concentrations and composition of biomass burning aerosol as before, the fraction of solar energy (between 300 and 700 nm) absorbed by organic compounds (HULIS) is 8.5 and 11.5% for  $\alpha=6$  and 7, respectively. In the 300–1200 nm range, the corresponding values are 6.4 and 8.6%. It is noteworthy that, even if we take 4 as the most conservative estimate of the Ångström exponent of the HULIS, the contribution to the absorption is still about 5% in the visible range, and about 3.9% in entire tropospherically relevant solar spectrum (300–1200 nm).

#### 4. Conclusions

The absorption properties of the humic-like substances (HULIS) isolated from biomass burning aerosols were similar in day and nighttime samples. While the single scattering albedo of pure HULIS particles of about 50–100 nm diameter is high, about 0.98 at 532 nm, they still absorb light rather effectively in the UV and visible range. Hence, due to this strong spectral dependence of light absorption by HULIS and due to the elevated amount of HULIS in biomass burning aerosol, a significant fraction of the light is absorbed by HULIS at shorter wavelengths. For instance, during the biomass burning period in the LBA-SMOCC campaign, up to 50% of the light absorption (at 300 nm) by aerosols in our study region in the Amazon basin is caused by (water-soluble) HULIS. Since the intensity maximum of the tropospherically relevant solar spectrum is at higher wavelengths, however, the contribution of the HULIS to the absorption over the entire

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spectrum is about 7%. Consequently the humic-like substances could have an active role in the radiative transfer and in photochemical processes. The absorption of HULIS in the UV range can contribute to the up to now unaccounted fraction of the absorption observed in the troposphere.

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**Table 1.** Optical (at 532 nm) and physical parameters of HULIS isolated from biomass burning aerosol. Assuming 5% uncertainty in the measurement of the absorption and scattering coefficient and 10% uncertainty in the determination of the number concentration, the uncertainty of the refractive index is within 3%. The intervals in the table show the confidence intervals of the parameters.

	Index of refraction, real part	Index of refraction, imaginary part	Mass absorp- tion coeff. ( $\text{m}^2 \text{g}^{-1}$ )	Mass scatter- ing coeff. ( $\text{m}^2 \text{g}^{-1}$ )	Density ( $\text{g cm}^{-3}$ )
HULIS – day	$1.653 \pm 0.013$	$0.00187 \pm 0.00016 \text{ i}$	$0.031 \pm 0.003$	$1.96 \pm 0.14$	$1.569 \pm 0.031$
HULIS – night	$1.685 \pm 0.010$	$0.00163 \pm 0 \text{ i}$	$0.029 \pm 0.001$	$2.53 \pm 0.13$	$1.502 \pm 0.018$

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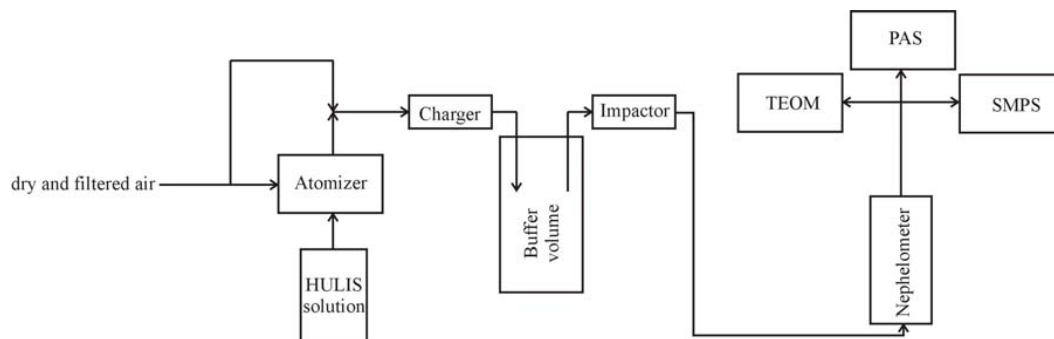
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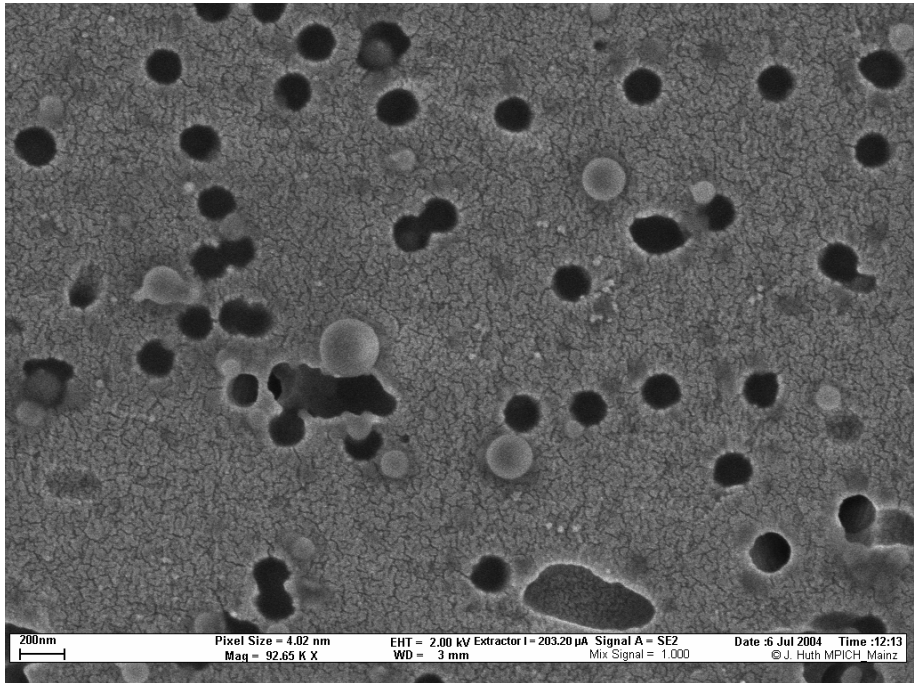
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**Fig. 1.** Experimental setup to measure the optical properties of HULIS.[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

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**Fig. 2.** SEM picture of particles generated by the TSI atomizer from a homogeneous solution of HULIS, which was isolated from the fine fraction of biomass burning aerosol from the Amazon basin.

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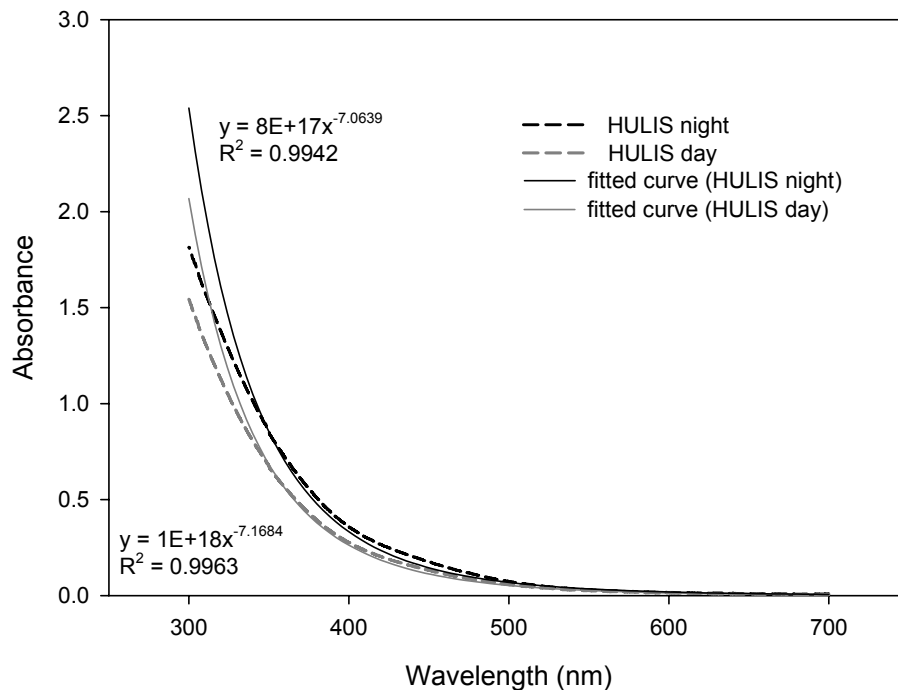
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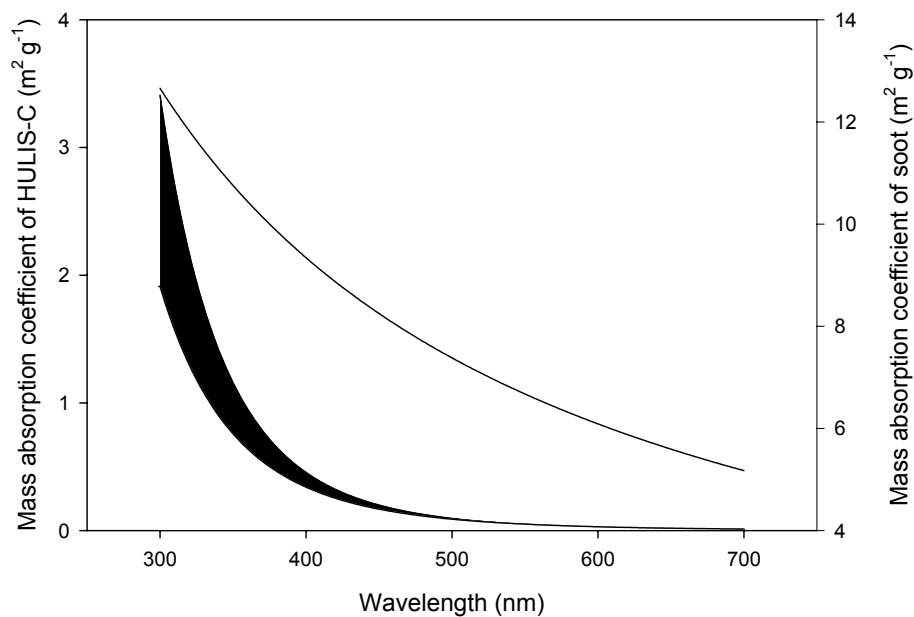
**Fig. 3.** The absorption spectra of HULIS isolated from day and night aerosol samples.

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**Fig. 4.** The estimated wavelength-dependence of the (carbon) mass absorption coefficient of soot (upper line) and HULIS (lower trace:  $\lambda=6$ , upper trace  $\lambda=7$ ).

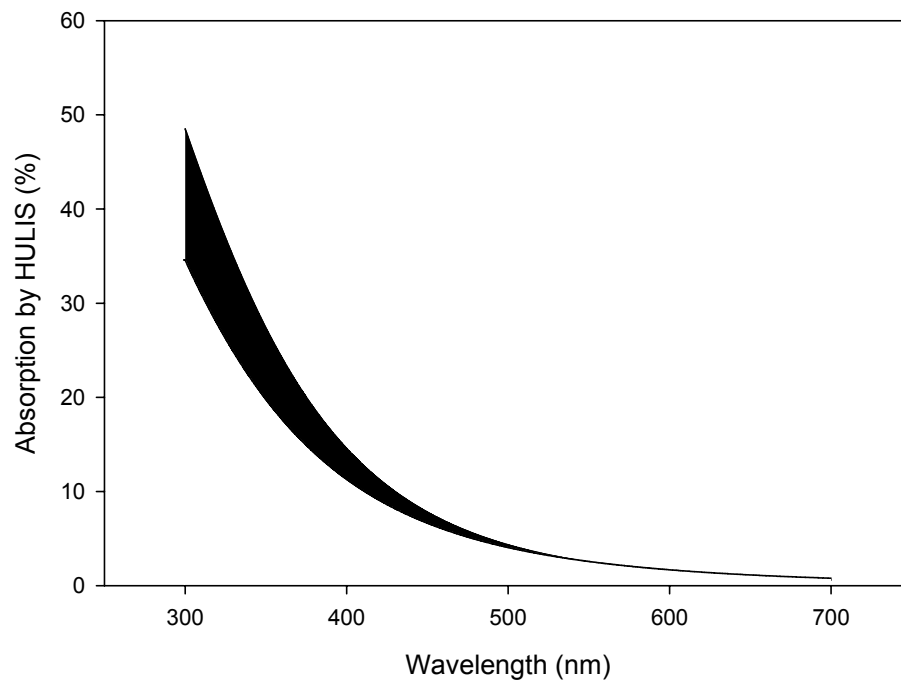
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**Fig. 5.** Estimated relative contribution of HULIS to light absorption in biomass burning aerosol (lower trace:  $\lambda=6$ , upper trace  $\lambda=7$ ).

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